

# Will ET Write or Radiate

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- **Mobility**
  - Not anathema  $\Rightarrow$  it helps!
  - Delay tolerant?  $\Rightarrow$  transmit when near base!
- **Interference Avoidance, Pricing & Spectrum Management**
  - Interference hurts  $\Rightarrow$  deal with it!

## 10 Years of Wireless Research

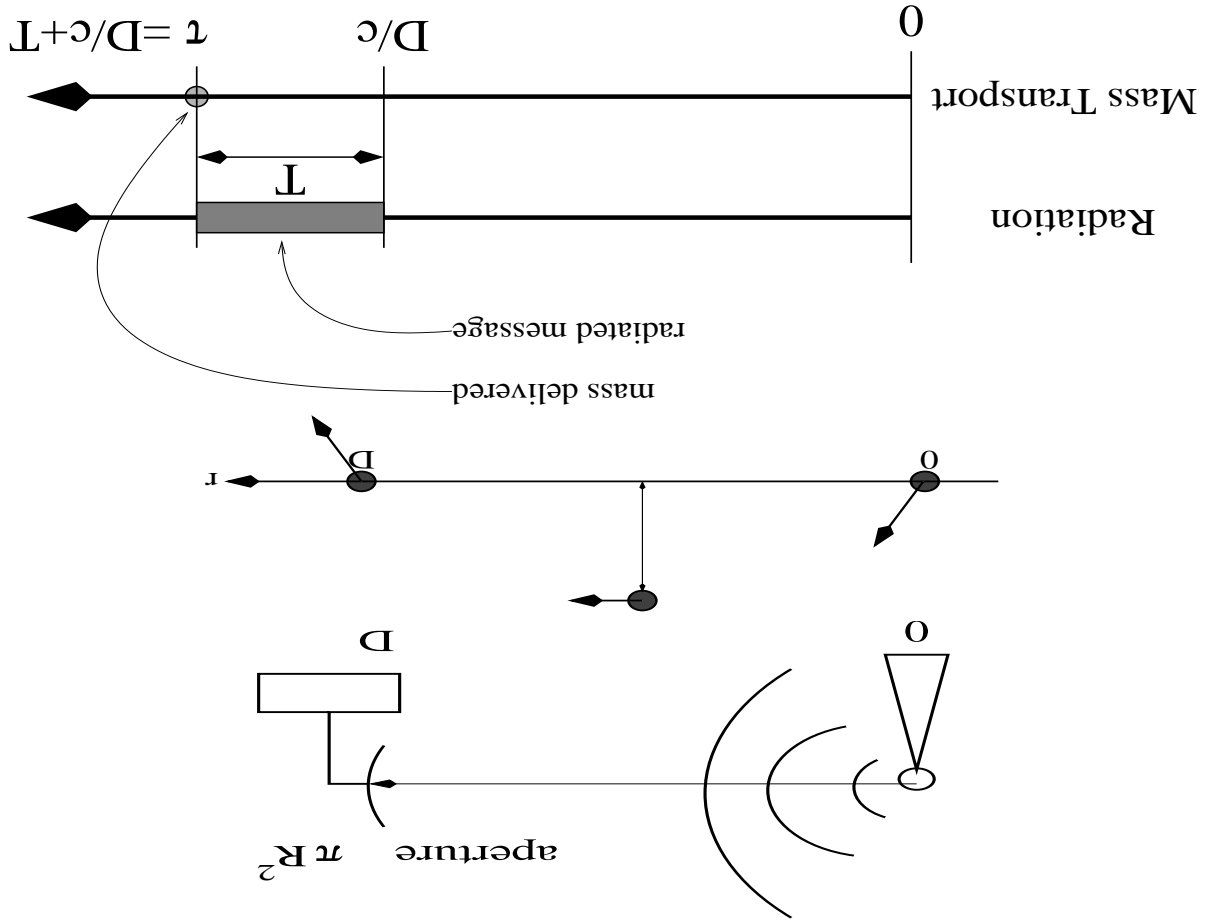
## A Hoary Old Epiphany

- RF interference is bad
  - Mobility is good
    - Channel really good when nearby
    - Can often tolerate delay
  - Storage density is increasing (faster than Moore)
    - **Go Postal!**
    - Forget RF! Write message down! Toss it to recipient!
    - Netflix! FedEx!
- (right???)

## A Little Empirical Rigor

- **Magnetic Storage with FeO<sub>2</sub>:  $2 \times 10^{17}$  bits/kg**
- **Optical Lithography with SiO<sub>2</sub>:  $3.85 \times 10^{18}$  bits/kg**
- **E-beam Lithography with SiO<sub>2</sub>:  $1.54 \times 10^{21}$  bits/kg**
- **STM with Xe on Ni:  $1.74 \times 10^{22}$  bits/kg**
- **RNA:  $3.6 \times 10^{24}$  bits/kg**

# A Little Analytic Rigor



Write or Radiate

## Radiation Energy Requirements

- Energy capture

$$v(D) = \frac{4\pi D^2}{AG}$$

- Bits a la Shannon:

$$B = TC = TW \log_2 \left( \frac{P}{GA} \frac{4\pi D^2}{N_0 W} + 1 \right)$$

- $E_r = PT$ ,

$$E_r = BN_0 \frac{4\pi D^2}{TW} \frac{AG}{B} \left[ 2^{\frac{TW}{B}} - 1 \right]$$

- Large  $TW$ :

$$E_r \geq BN_0 \frac{4\pi D^2}{AG} \ln 2$$

## Some (free)space) Rocket Science

- Average velocity

$$\frac{1}{D} \int_0^{\tau} v(t) dt = \bar{v} = \frac{D}{\tau} [E[v(t)]]$$

- Minimum imparted energy subject to  $\bar{v} = \frac{D}{\tau}$ :

$$E_* = \min_{v(t)} \max_t h(v(t))$$

- Jensen says

$$E_* = \min_{v(t)} \max_t h(v(t)) \geq \min_{v(t)} E[h(v(t))] \geq h(\bar{v})$$

with equality iff  $v(t)$  is constant =  $\bar{v}$   
(Jensen's Inequality Rock(et)s!)

## Minimum Transport Energy

- GIVEN:  $h(\nu)$  and  $\nu$

$$E_* = h(\nu)$$

- $h(\nu) = mc^2 \left( \frac{1}{1 - \frac{\nu^2}{c^2}} - 1 \right)$  :

$$E_* = mc^2 \left( \frac{1}{1 - \left(\frac{\nu}{c}\right)^2} - 1 \right)$$

- $h(\nu) \approx \frac{1}{2}m\nu^2$  :

$$E_* \approx \frac{1}{2}m\nu^2$$



## But What About Gravity?

- $q(x)$  potential energy:

$$\mathcal{E}(t) = h(v(t)) + q(x(t))$$

- Energy minimization:

$$E_* = \min_x \max_t \mathcal{E}(t, x(t), v(t)) \geq \min_x \int_0^1 \mathcal{E}(t, x(t), v(t)) dt$$

- Calculus of variations:

$$0 = \frac{d}{dt} \left( \frac{\partial \mathcal{E}}{\partial v} \right) - \frac{\partial \mathcal{E}}{\partial x}$$

$$0 = \ddot{x}h''(x) - q'(x)$$

## Potential Field Results

- Non-relativistic:
- $m\ddot{x} = q'(x)$
- $q'(x)$  is force at position  $x$ :  $\rightarrow$  “free fall”
- Freefall?  $\rightarrow \mathcal{E}(t) = \text{constant}$
- $\mathcal{E}(t)$  constant  $\rightarrow$  minimization satisfied with equality, so ...



- Low speed:
- $m\ddot{x} = q'(x)$
- $q'(x)$  is force at position  $x$ :  $\rightarrow$  “free fall”
- Freefall?  $\rightarrow \mathcal{E}(t) = \text{constant}$
- $\mathcal{E}(t)$  constant  $\rightarrow$  minimization satisfied with equality, so ...

## Potential Fields Results

## Gravity Summary

**Artillery:** adds a factor of 2 to energy

**Escape:** small penalty if  $v > 2 \times$  escape velocity

## Inscribed Matter Energy Requirements

- Message size  $B$ , mass information density  $\tilde{\rho}$

$$E_w \approx \frac{1}{2} \frac{B}{\tilde{\rho}}^{-2} = \frac{1}{2} \frac{B}{\tilde{\rho}} \left( \frac{\delta}{c} \right)^2$$

– Brisk Walk:  $\delta = 3 \times 10^8$

– Jetliner:  $\delta = 10^6$

– Earth Escape:  $\delta < 2.7 \times 10^4$

– Solar Escape:  $\delta > 7.1 \times 10^3$

– Milky Way Escape:  $\delta > 6 \times 10^2$

## Radiation to Transport Energy Ratio

- Definition:

$$\Omega = \frac{E_r}{E_w}$$

- Large  $TW$ ,  $\delta \gg 1$

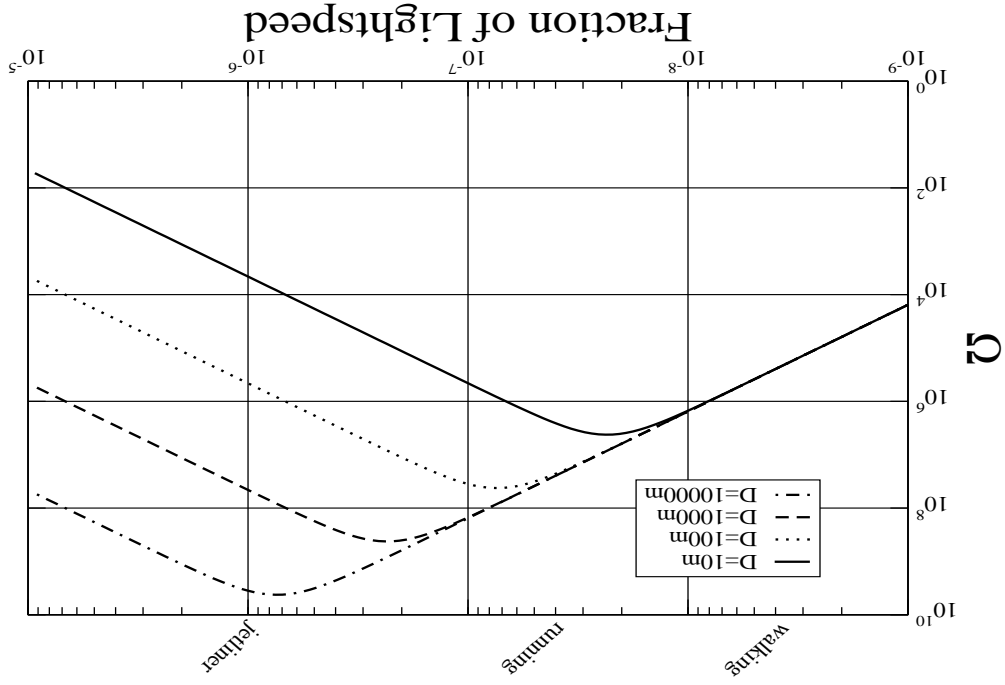
$$\Omega \geq \left[ \frac{\tilde{p}_{N_0}}{c^2} \right] \left[ \frac{4\pi D^2}{AG} \right] (2 \ln 2) \delta^2$$

$$\begin{aligned} \text{Normalized Aperture} &\equiv \mathcal{A} = \frac{\lambda}{2R} \\ \text{Normalized Distance} &\equiv \mathcal{D} = \frac{D}{2R} \end{aligned}$$

$$\Omega \geq \left[ \frac{\tilde{p}_{N_0}}{c^2} \right] \left[ \frac{8}{\pi^2} \left( \frac{\mathcal{D}}{\mathcal{A}} \right)^2 \right] (2 \ln 2) \delta^2$$

# Isotropic Radiation vs. Terrestrial Artillery

$$p = 3 \times 10^{24}, R = 5\text{cm}, \text{Temperature } 300\text{K}$$



## Chip to Chip Laser Links

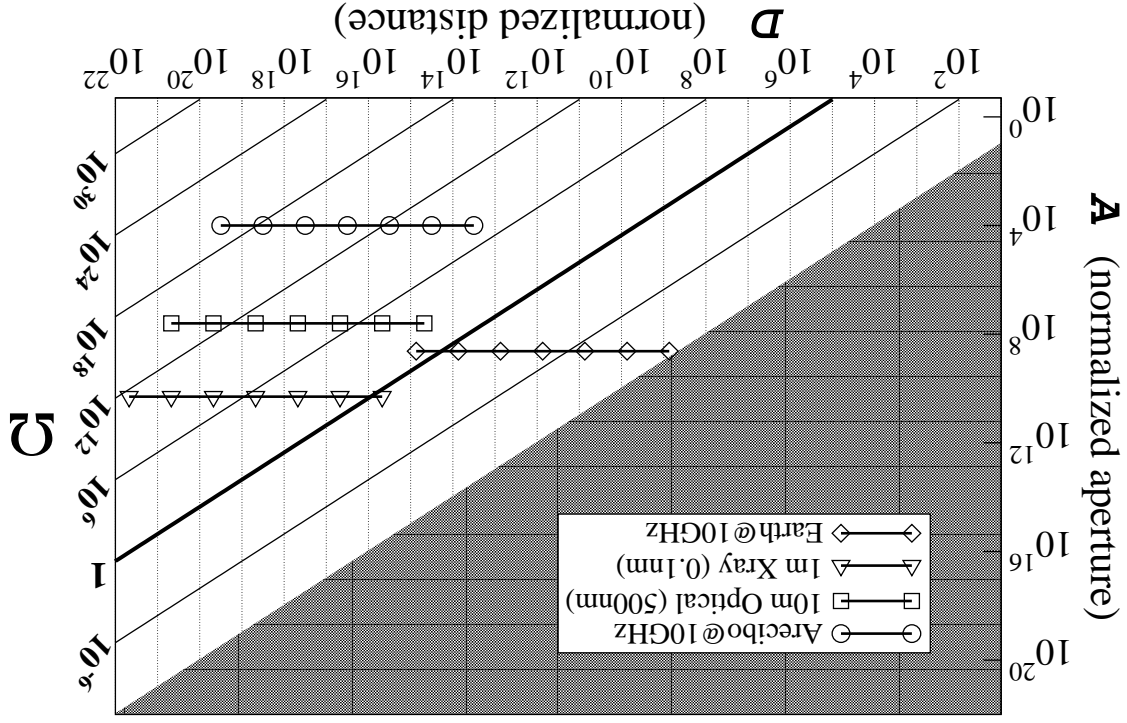
- Temperature: 300K
- Delay:  $\delta = 10^9$
- Aperture:  $R = 5\mu\text{m}$
- Wavelength:  $\lambda = 1\mu\text{m}$
- Distance:  $D = 10\text{cm}$
- $\Omega = 5.17 \times 10^{-14}\text{p}$

**MUCH better to toss magnetic chips!**



# Interstellar

$p = 10^{22}$ ,  $\delta = 10^3$ , Temperature 3K



(from Nature 431, pp.47-49, September 2, 2004)

## Other Issues

### General

- Radiation is broadcast – IM still wins (at reasonable receiver densities)
- Inscription energy budget – generally small

### Interstellar

- Radiation needs repetition – penalty can be very large
- Matter deceleration at target – not terrible for  $\delta \gtrsim 10^3$
- Cosmic ray bombardment – even with shielding IM wins
- Advertisement at target – still studying

## No, inscribed matter still wins!

- Radiation illuminates many  $\rightarrow$  matter penalty
- Milky Way stellar density  $2.8 \times 10^{-2}$  stars (LY) $^{-3}$
- Spherical galaxy, isotropic radiation, Arecibo-Arecibo
- $R = 10^4$  LY:  $1.13 \times 10^{11}$  stars (but  $\Omega = 10^{28}$ )
- $R = 10^6$  LY:  $1.13 \times 10^{17}$  stars (but  $\Omega = 10^{32}$ )

**Is Radiation Better for Broadcast?**

- **Matter Inscription/Readout Energy and Time**
  - Can be reversible and arbitrarily fast (R. Landauer)
- **Empirical energy calc:**
  - 60000 ATP/second for 20 minutes: 4639 Kbase of E-coli
  - $8 \times 10^{-20}$  J per ATP molecule
  - $6.2 \times 10^{-19}$  J bit<sup>-1</sup>.
  - $E^*$  at earth escape:  $1.68 \times 10^{-17}$  J bit<sup>-1</sup>.
- **Construction energy probably not a problem**

**Does Inscription Energy/Speed Eat Budget?**

## Radiation Needs Repetition

- Civilization Birth Rate:  $\alpha = 1/10^9$  per year
- Civilization Extinction Rate:  $\beta = 1/10^6$  per year
- Success criterion  $0 \leq \Phi \leq 1$
- Now many repetitions  $N$  (optimally placed)?
- $\Phi = 0.99 \rightarrow N = 2 \times 10^5$
- $\Phi = 0.9999 \rightarrow N = 2 \times 10^7$ ,

## Package Deceleration Penalties

- Need exhaust braking
- Energy penalty (excess mass):  $e \frac{\delta g I_{sp}}{c}$
- $I_{sp} \equiv$  Specific Impulse
  - Chemical:  $10^2$
  - Nuclear Electric:  $10^4$
  - Fusion:  $10^6$
- $I_{sp} = 20,000, \delta = 1000 \rightarrow$  penalty 4.6
- $\delta = 100$  or  $I_{sp} = 2000 \rightarrow$  penalty  $4.4 \times 10^6$

## Cosmic Insults

- **Insults:**
  - High energy particle bombardment
  - Heating (diffusion)
  - Ion tracks, dislocations, subatomic cascades
- **Shielding:**
  - 10 million years at 10% bacteria viability: 3 m radius rock
  - ( $3\text{g cm}^{-3}$  density)
  - $3.4 \times 10^6$  penalty
- **Clever Composition, Coding and Correction:**
  - ??????????

## Advertisement Methods

- Big rock?
- Dust?
- Embedded dust & rock (comet)?
- Active Probe (Bracewell)?
- Colonization via panspermia (Crick)?



## Paradigm-Piercing Punchlines

- Incribed matter messaging is **NOT** ridiculous
  - FedEx and Netflix are on to something
- Finesses Gupta-Kumar *ad hoc* nets  $\sqrt{N}$  problem?
  - little data missiles
- Chip-to-chip or mote-to-mote (high latency) communication?
  - smart dust tossing inscribed dust
- **ET might write not radiate**
  - look for artifacts not radio

**Learn more:**

**Nature** cover story (431, pp.47–49, September 2, 2004)  
<http://www.winlab.rutgers.edu/~crose/cgi-bin/cosmic19.html>